

Liquids Discovered on Titan's Surface

Cassini's radar instrument was the first to confirm the presence of liquid lakes on Titan's surface – in the north polar region.



Titan's surface from Huygens DISR

- The Huygens Probe landed in a dark floodplain, amid rounded “rocks” of water ice
- The rounded nature of the water ice rocks show erosion due to tumbling down liquid CH₄ streambeds
- These are analogous to small pebbles found in rivers on Earth
- After landing, the probe’s batteries survived another 72 minutes



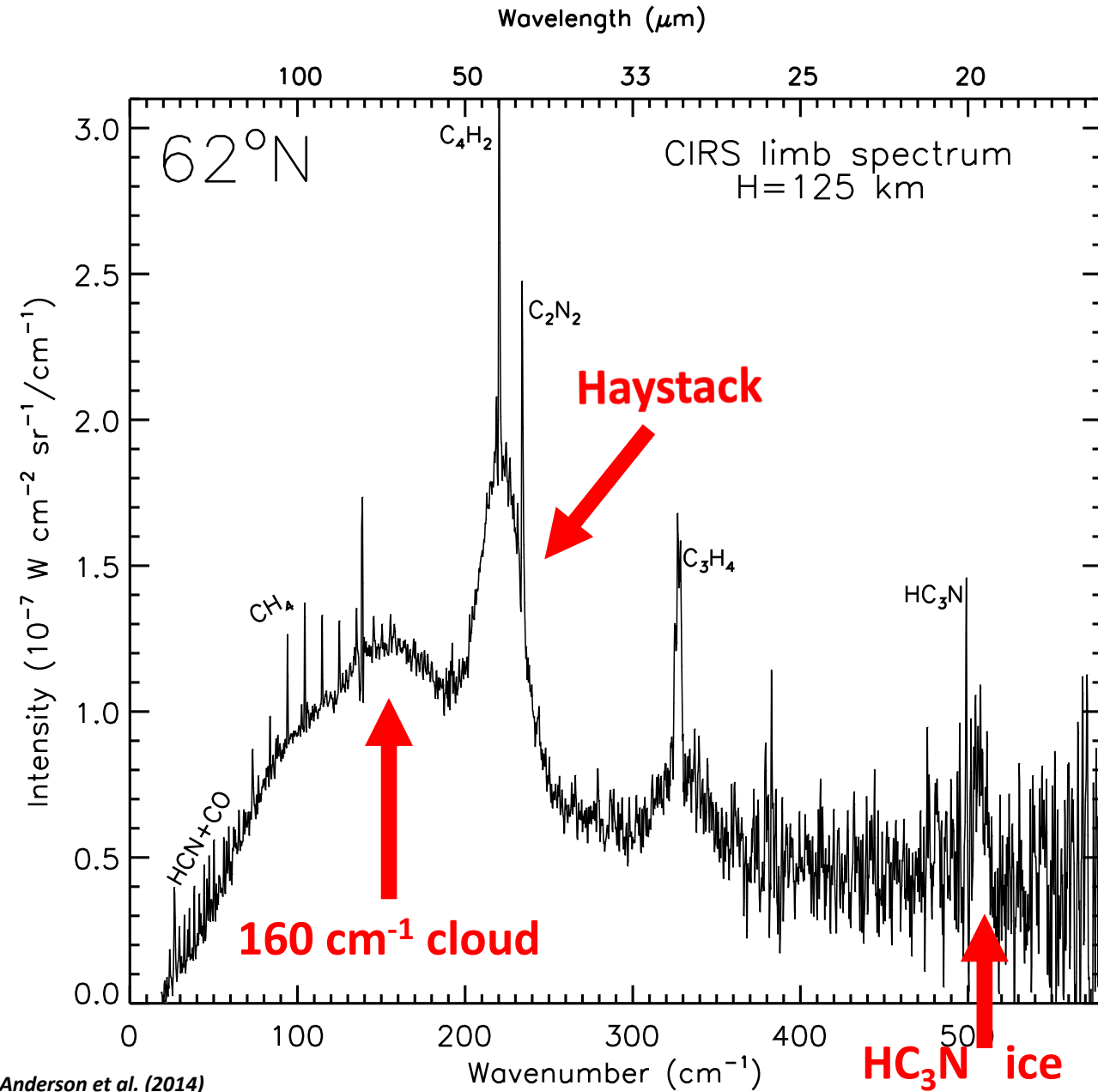
Titan's Stratospheric Clouds

- In Earth's stratosphere, clouds rarely form due to it being very dry
- Titan's stratospheric ice clouds are very different, and by far the most complex of any observed in the solar system
- Titan's stratospheric clouds form as a direct result of the general circulation pattern
- There are over a dozen organic vapors condensing out to form a suite of pure and co-condensed ices, typically observed at high winter polar latitudes
- Co-condensation processes commence when some of these organic vapors enter altitude regions where they can simultaneously saturate
- Most of these ices will diffuse throughout Titan's lower atmosphere and eventually precipitate to the surface, where they are expected to contribute significantly to Titan's regolith



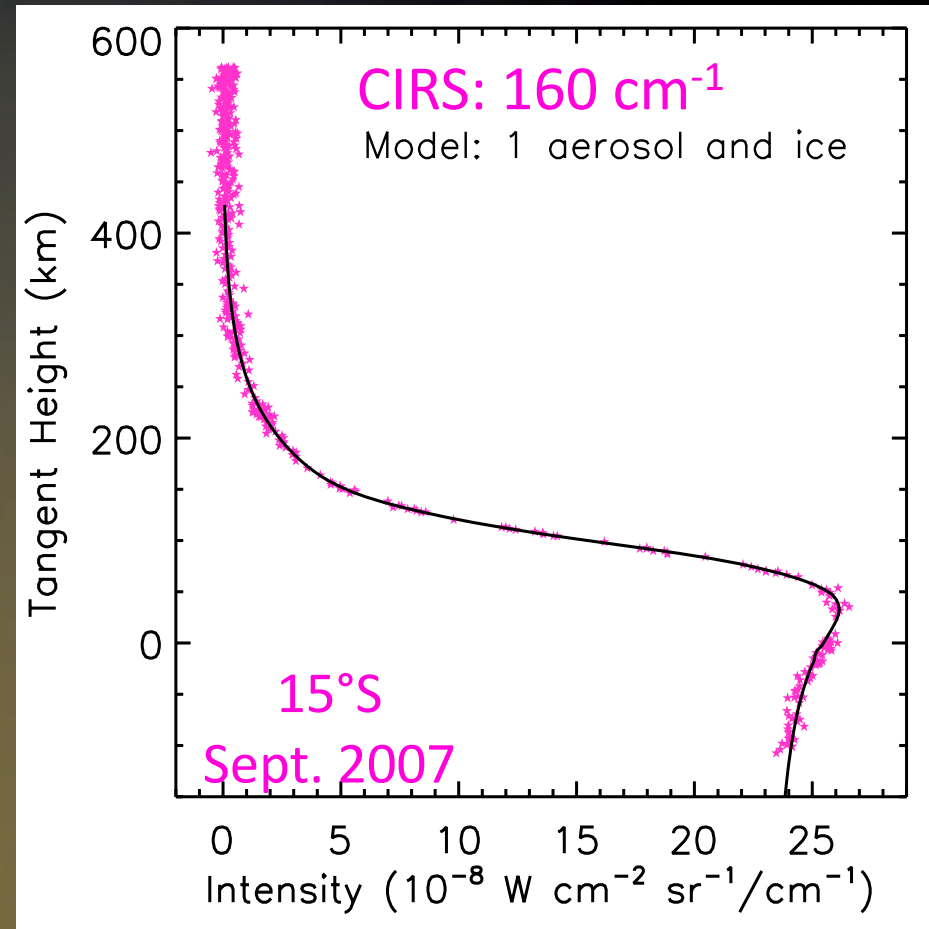
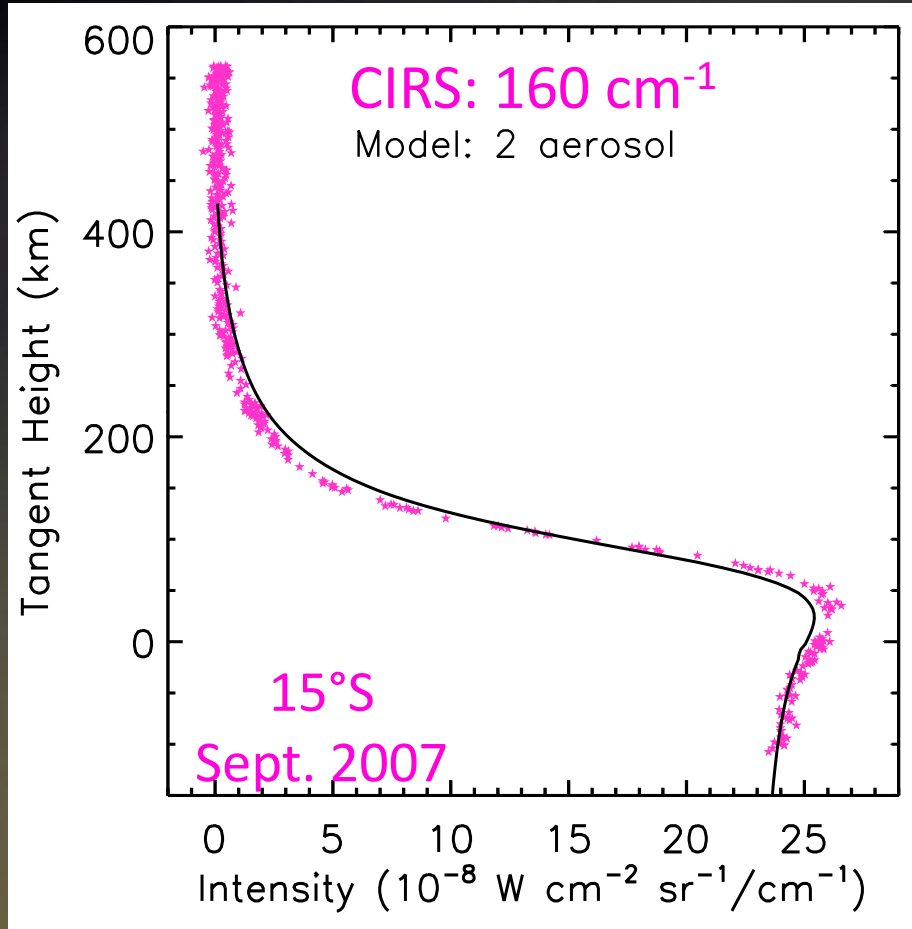
First Stratospheric Clouds Detected by CIRS

- Titan spectra recorded at 62°N in July 2006 with the Cassini CIRS far-IR FOV centered on ~ 125 km tangent height
- Season was northern winter
- Numerous gases and ices are visible
- The CIRS far-IR limb integration spectrum shows three chemically-different stratospheric ice clouds: co-condensed HCN/ HC_3N , the Haystack, and the ν_6 band of HC_3N

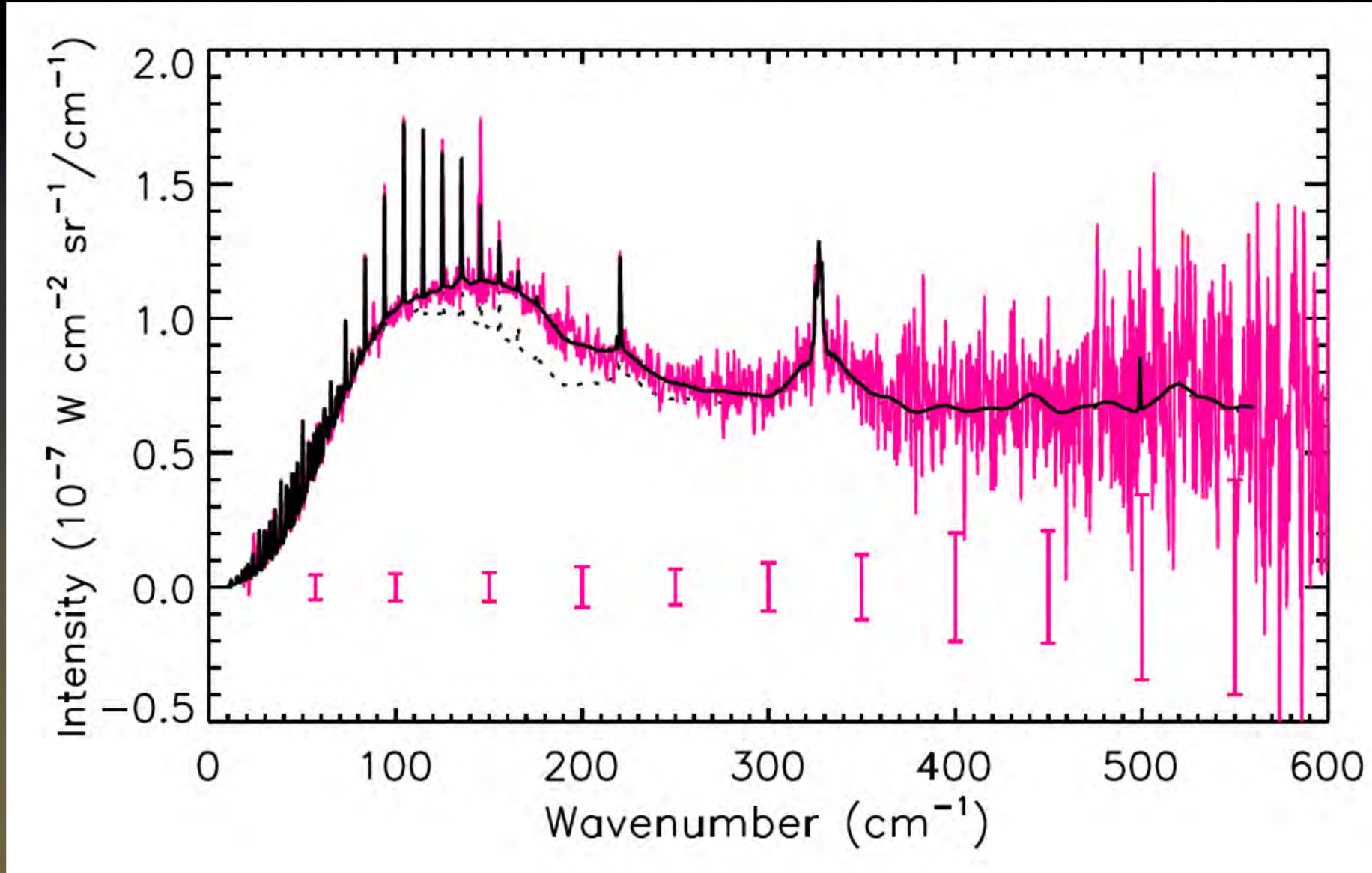


Discovery of co-condensed ice clouds

Based on the cloud top altitude and thickness, we already knew that the 160 cm^{-1} ice cloud should predominately contain HCN and HC_3N ices.

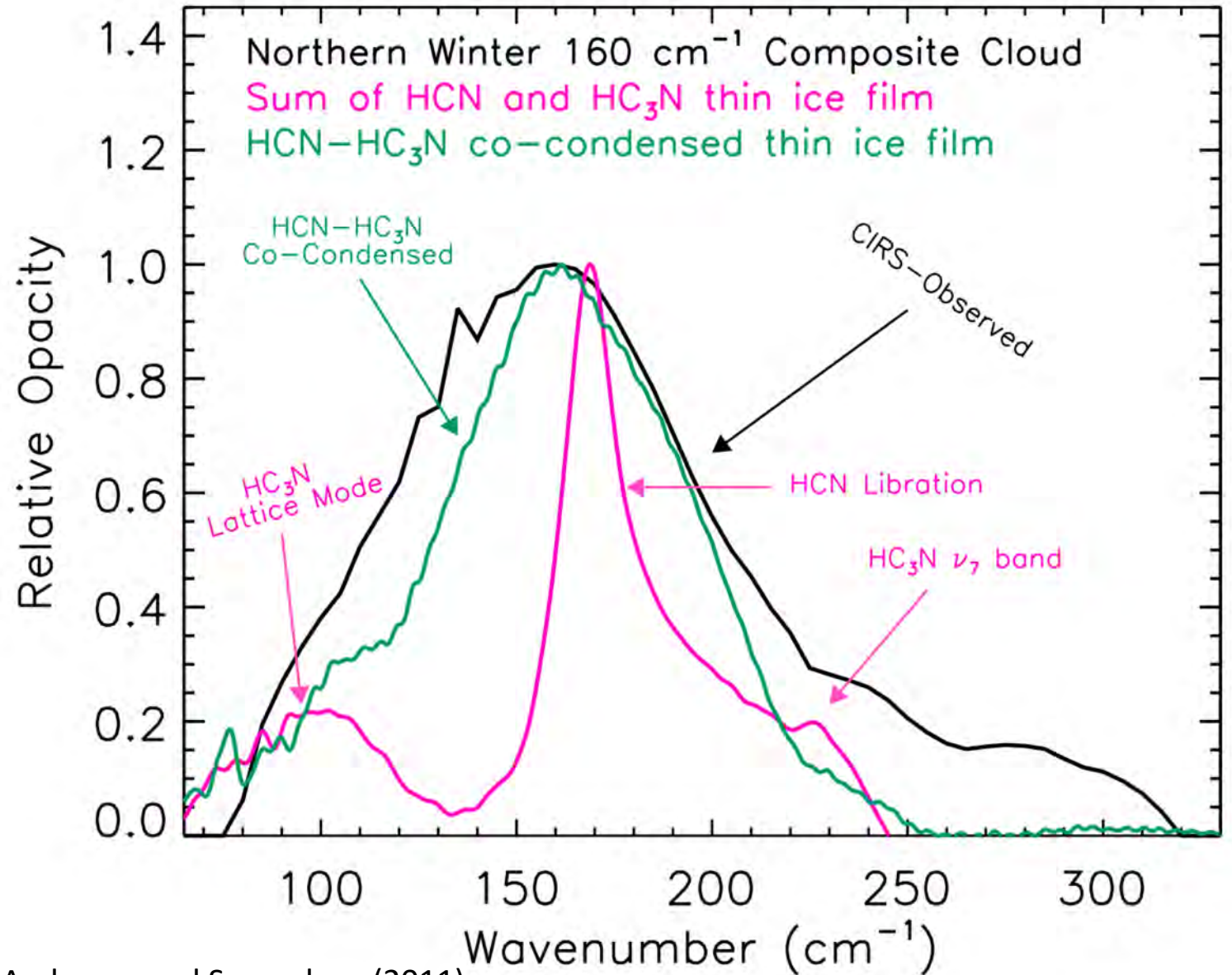


Discovery of 160 cm^{-1} ice cloud

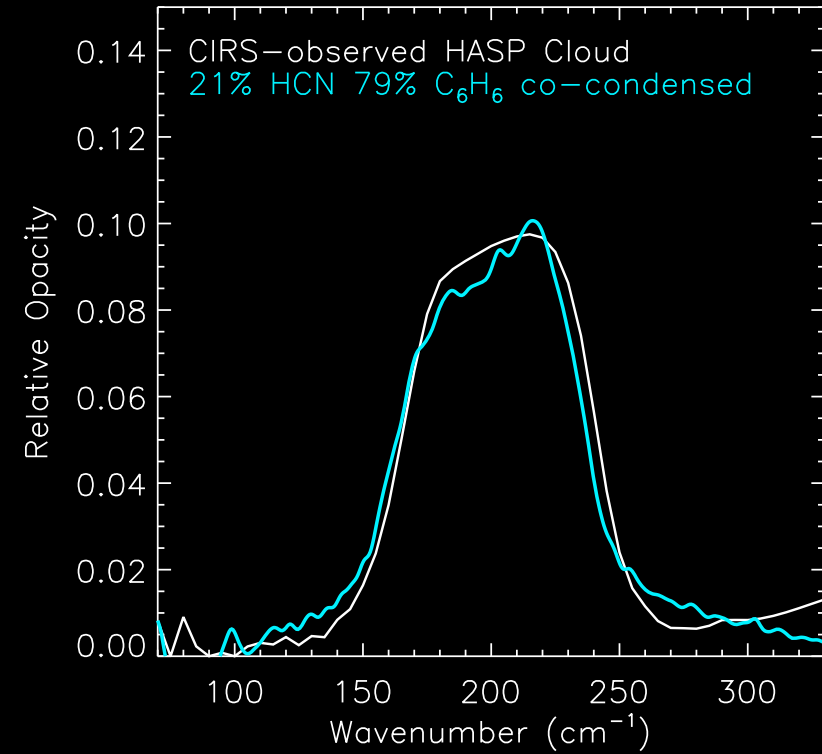
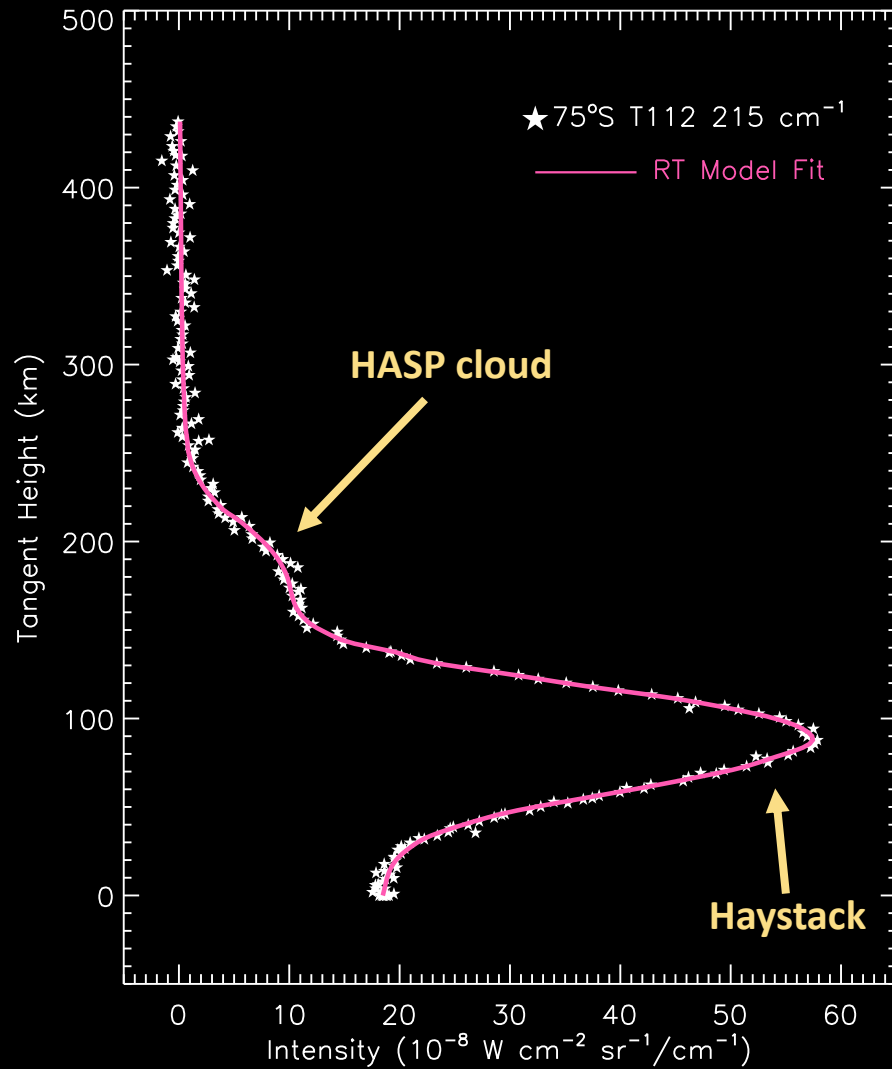


CIRS 160 cm^{-1} ice cloud

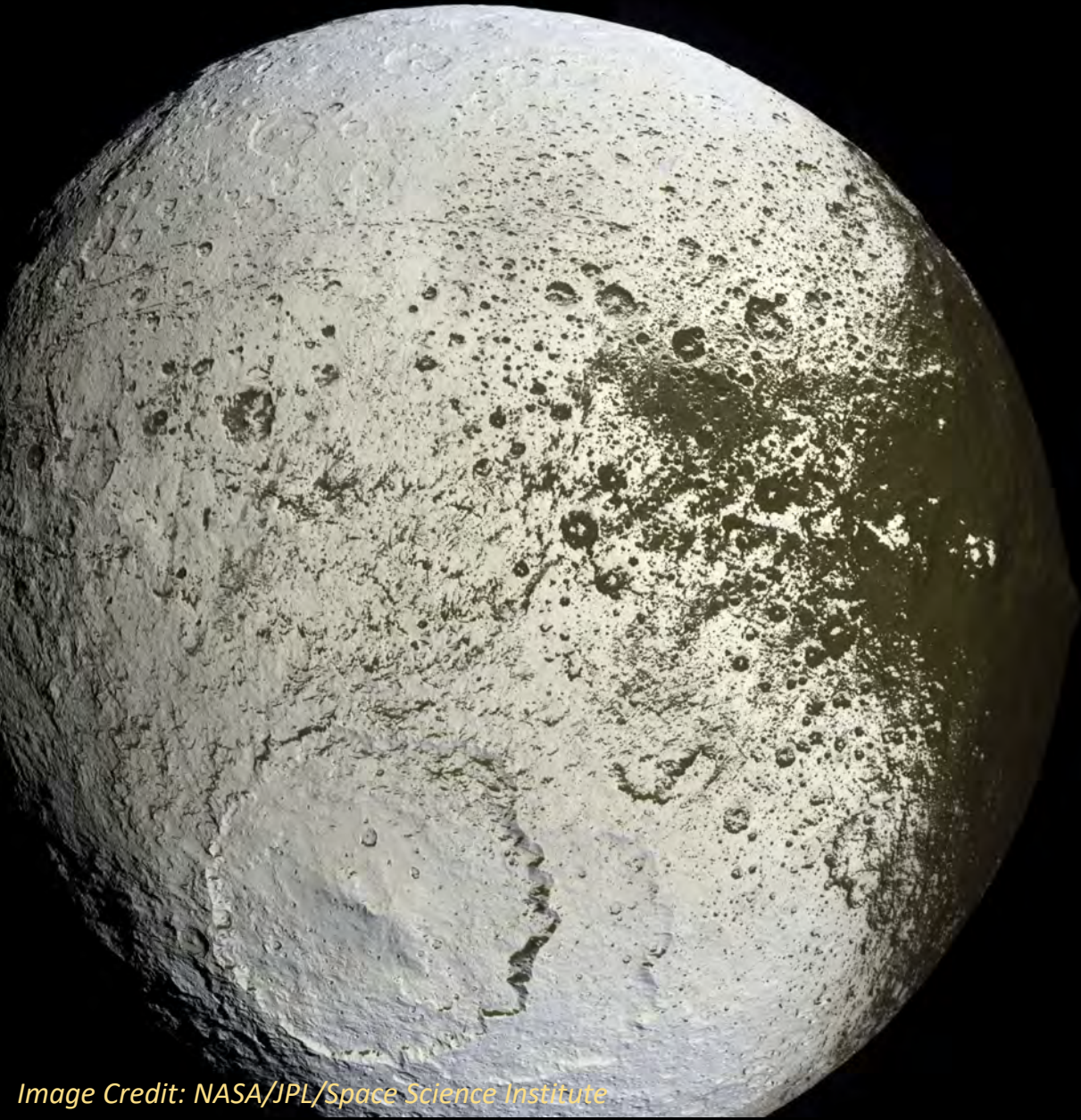
Thin ice films of layered HCN and HC_3N ice do not reproduce the CIRS-derived cloud opacity spectral dependence.



CIRS-discovery of another co-condensed ice cloud

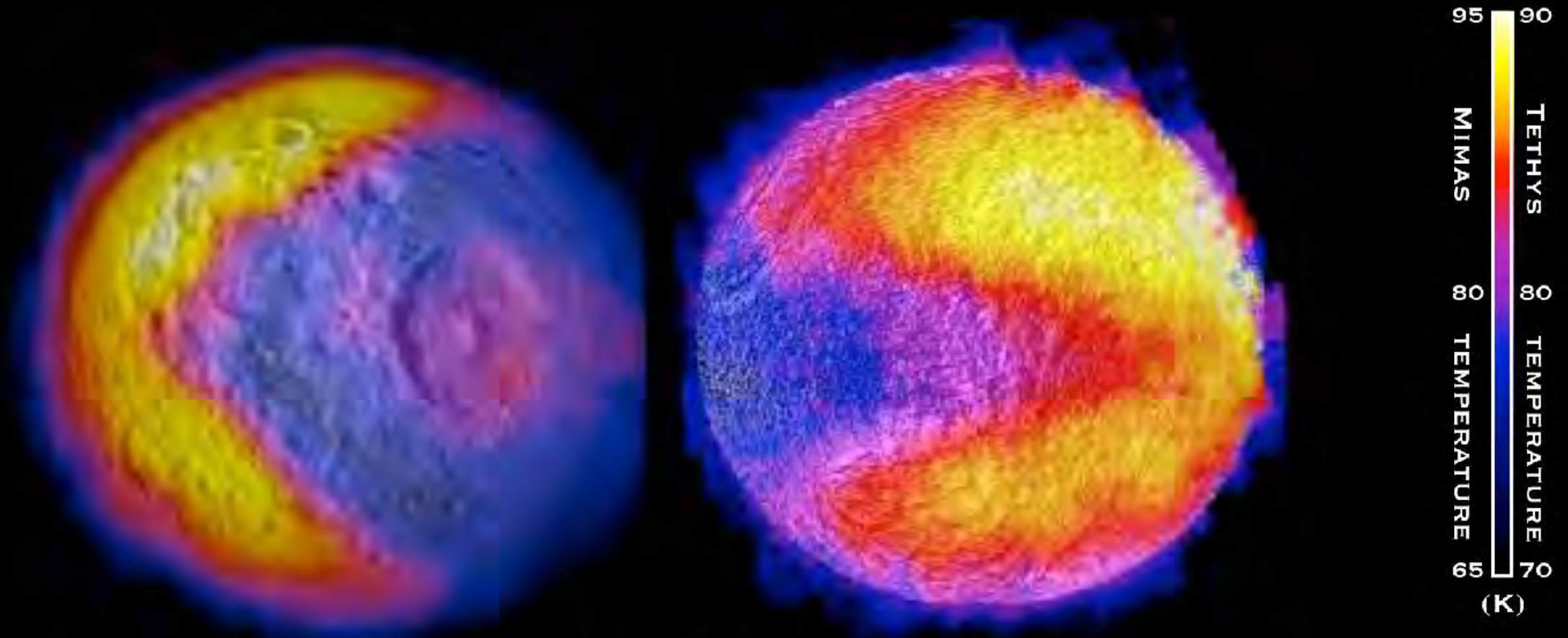


Iapetus



- Cassini observed the first high-spatial resolution image of Iapetus' trailing hemisphere - as bright as snow
- The bottom half shows a 280-mile wide impact basin. This is just 1 of 9 on the moon
- The leading hemisphere is as dark as coal, most likely covered in condensed nitrile ices, hydrated minerals, and other carbonaceous minerals
- The dark material may be due thermal segregation, a result of Iapetus' slow, 79-day rotation

Mimas, Tethys, and “Pac-Man”



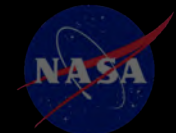
High spatial resolution temperature and image maps revealed unexpected hot regions that resembled “Pac-Man.”

Saturn



Cassini detected and tracked the progression of an early northern spring large storm in Saturn's northern hemisphere.

The storm extended ~9,000 miles from north-to-south, the largest observed on Saturn in the last 20 years.

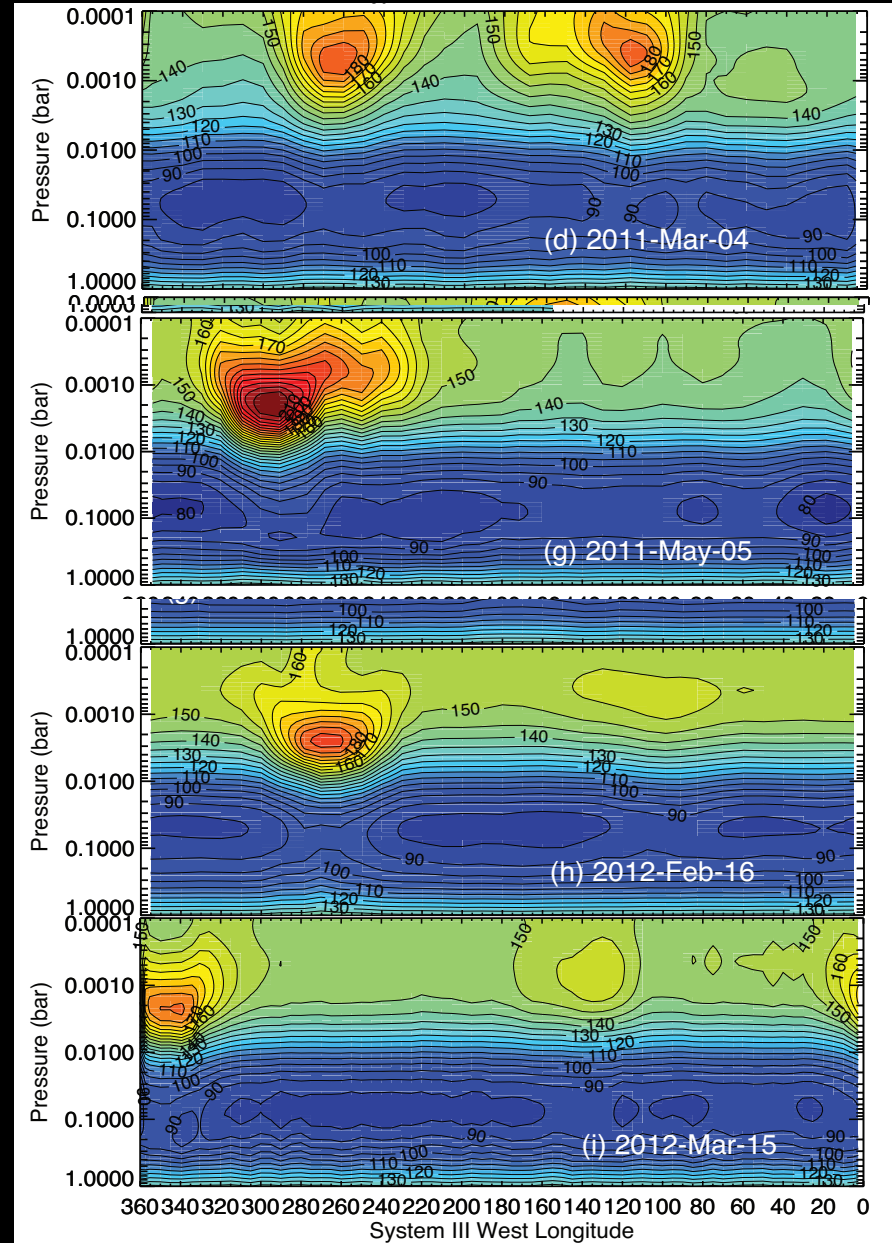


IR Hotspots in Saturn's Monster Storm

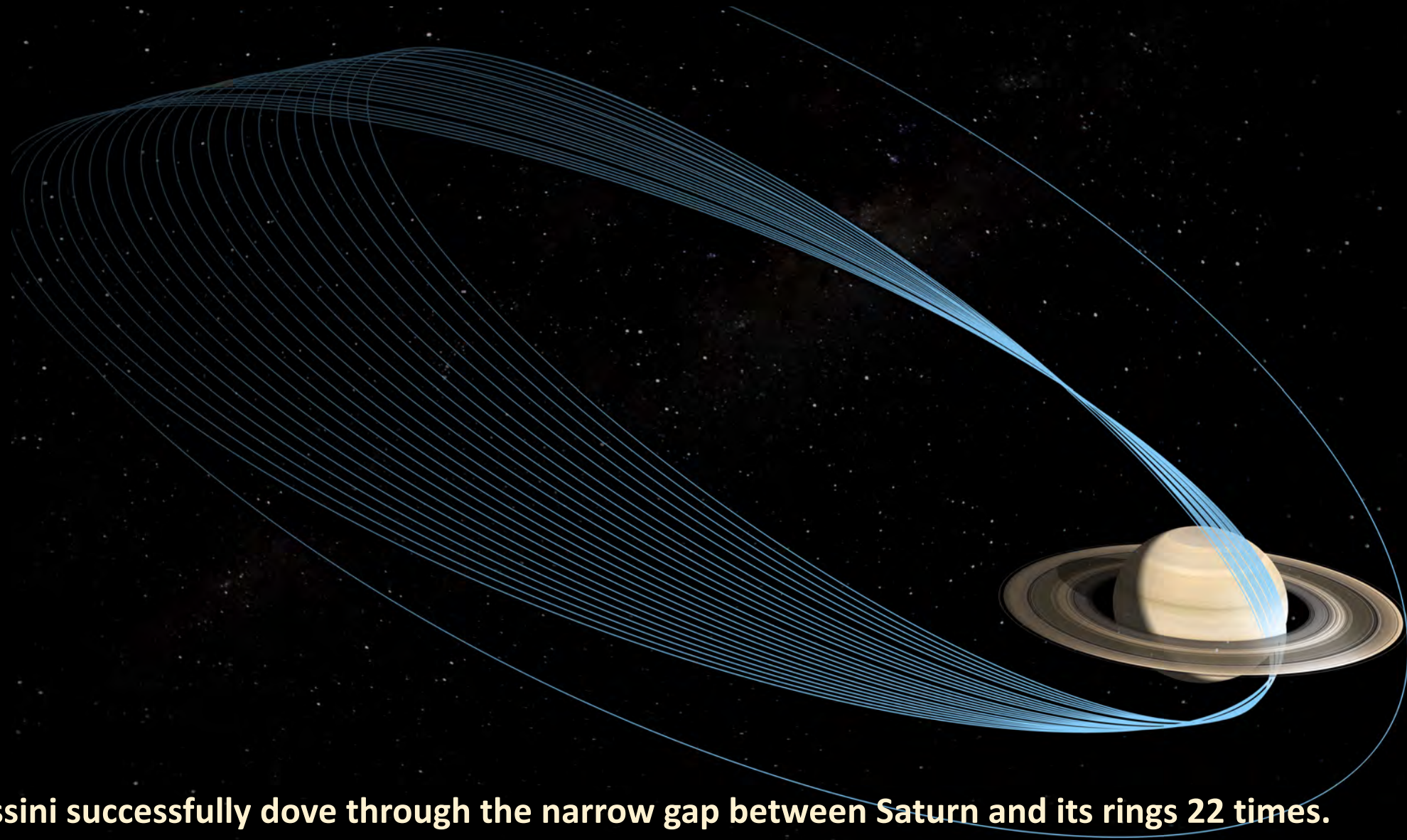
Over the next 2 years, when the storm was invisible to Cassini's camera, CIRS observed two vast beacons of hot air in Saturn's stratosphere.

As the tropospheric cloud formed, waves of heat travelled hundreds of miles upwards, depositing their energy as two vast "beacons" of hot air in Saturn's stratosphere.

Saturn's Monster Storm



The Grand Finale Begins – 26 April 2017

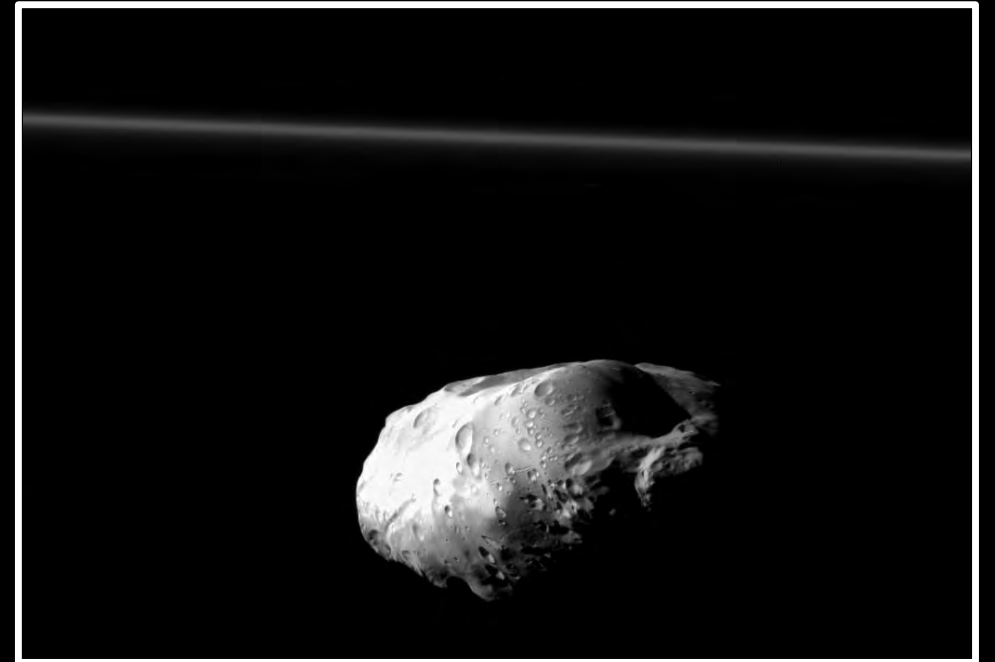
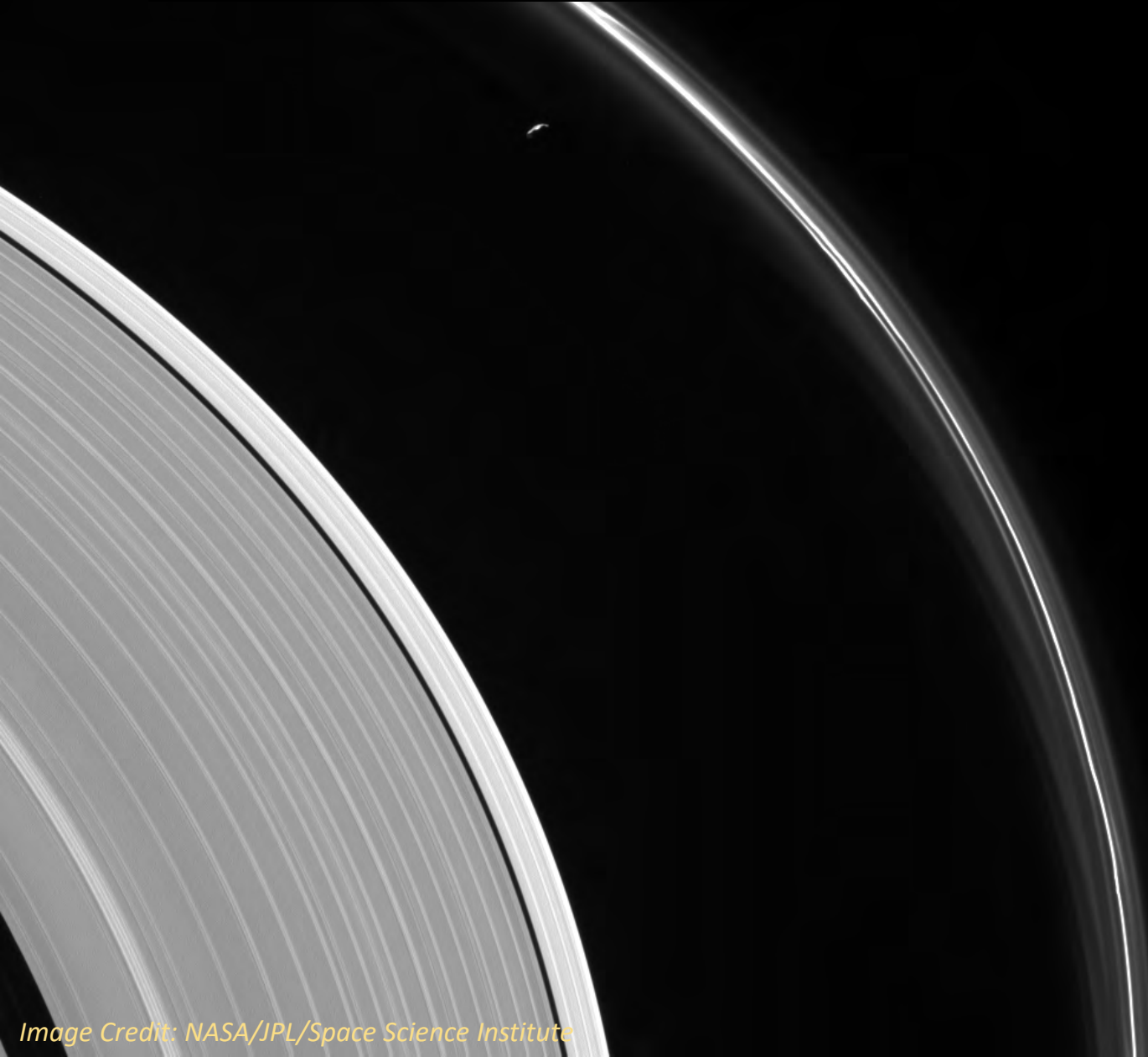


Cassini successfully dove through the narrow gap between Saturn and its rings 22 times.

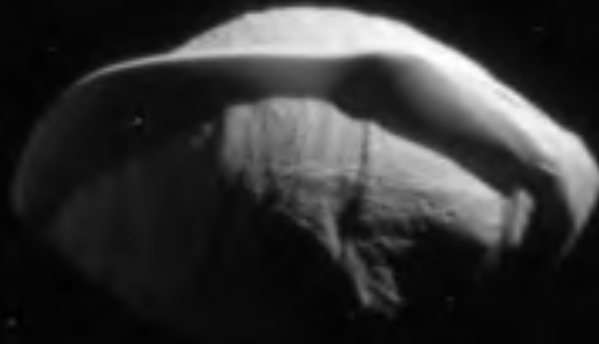
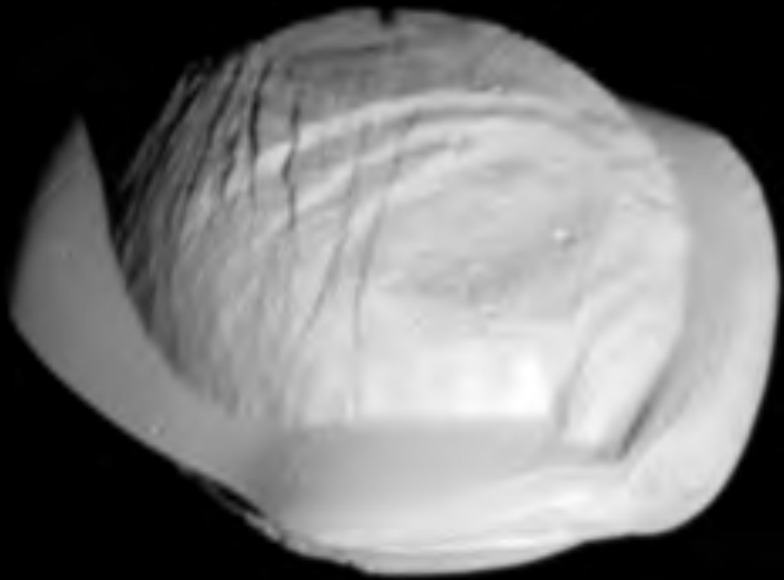
Farewell to Titan – 21 April 2017

- Throughout the entire mission, Cassini has used Titan's gravity for major trajectory changes in nearly all of its close flybys of the moon
- During T-126, Cassini passed within 600 miles of Titan and used the moon's gravity as a pivotal point to alter the shape of Cassini's orbit in a tragic way
- Even if we lost communication with Cassini, the spacecraft was now on an eventual collision course with Saturn
- T-126 marks the end of Cassini's Ring-Grazing Orbits, and sets the stage for the beginning of the mission's Grand Finale

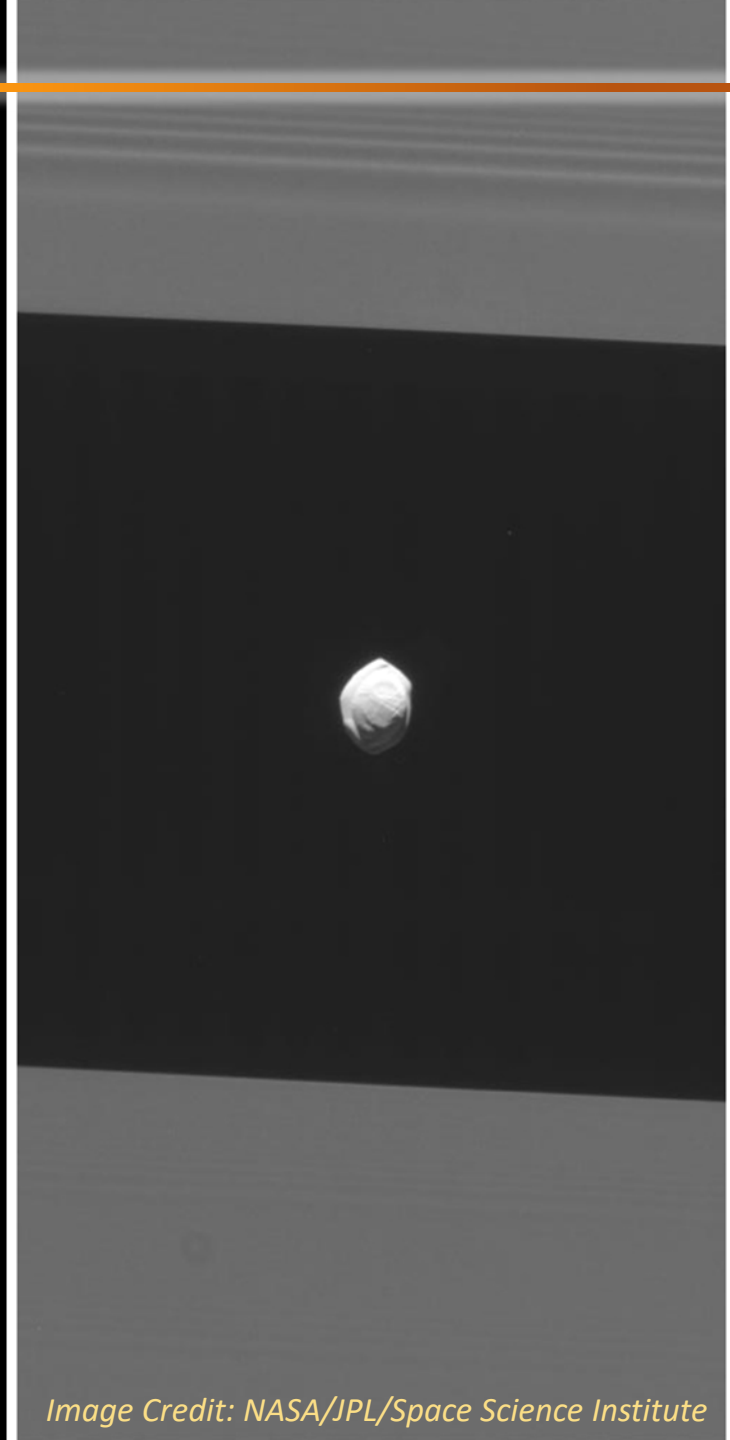
Prometheus and the F-ring



Pan Up Close



Pan is embedded in Saturn's A ring, causing a gap – the Encke gap – in the ring as it orbits.



Saturn Up Close

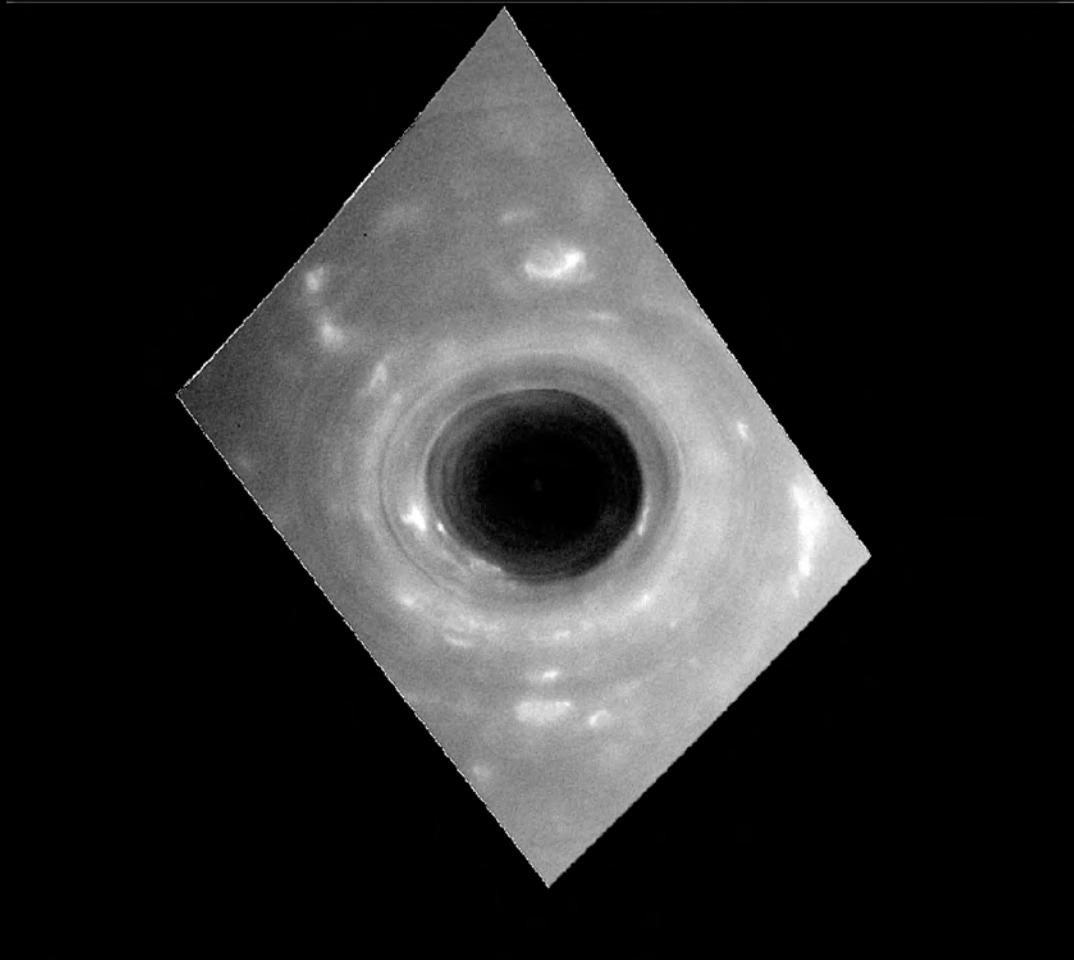
A close-up view of Saturn's north pole, showing a prominent hexagonal jet stream. The hexagon is a large, dark, swirling feature with a central dark spot. The surrounding atmosphere is lighter and shows some cloud patterns. The planet's rings are visible in the lower-left corner, appearing as a series of parallel lines.

Saturn's hexagon-shaped jet stream around the planet's north pole

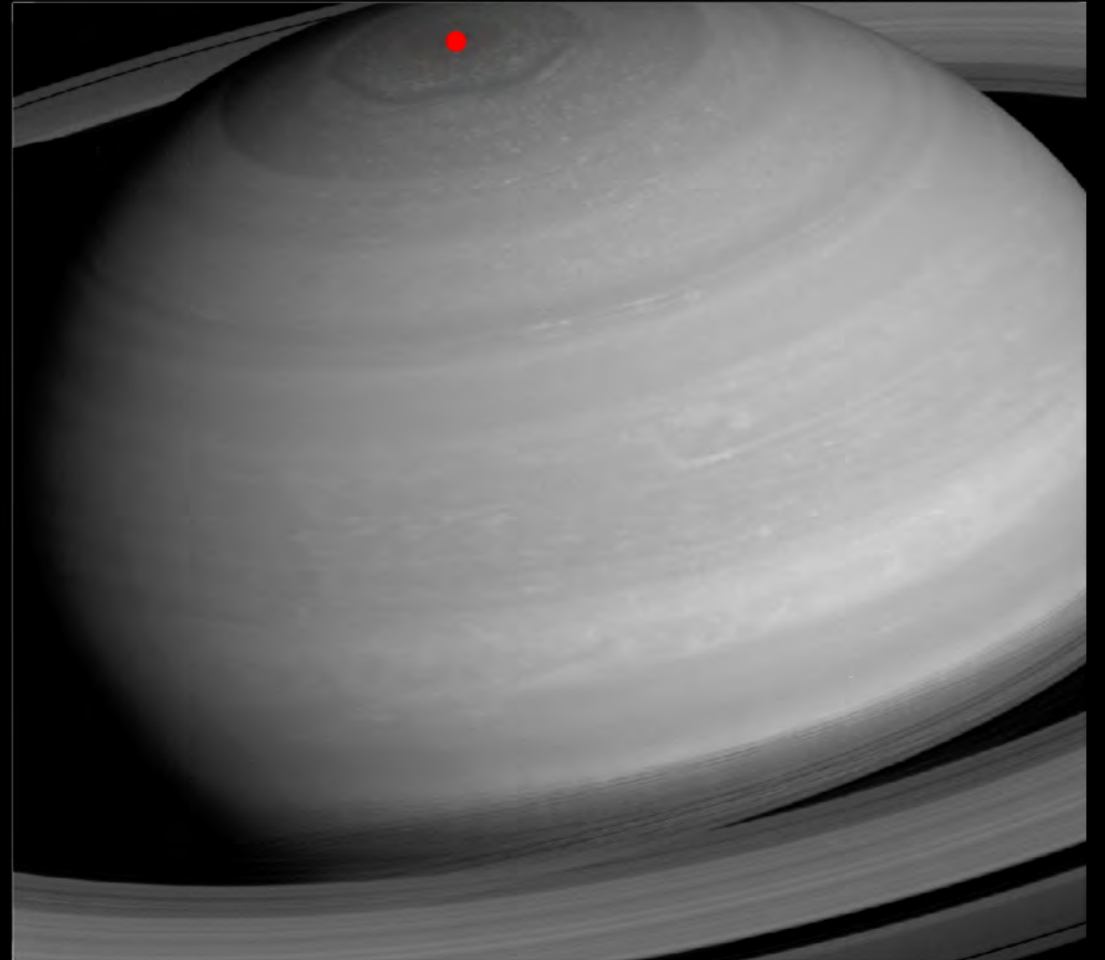
Each side of the hexagon is about as wide as the Earth.

First Dive between Saturn and its Rings

Camera View



Reference View



Last Enceladus Plume Observation

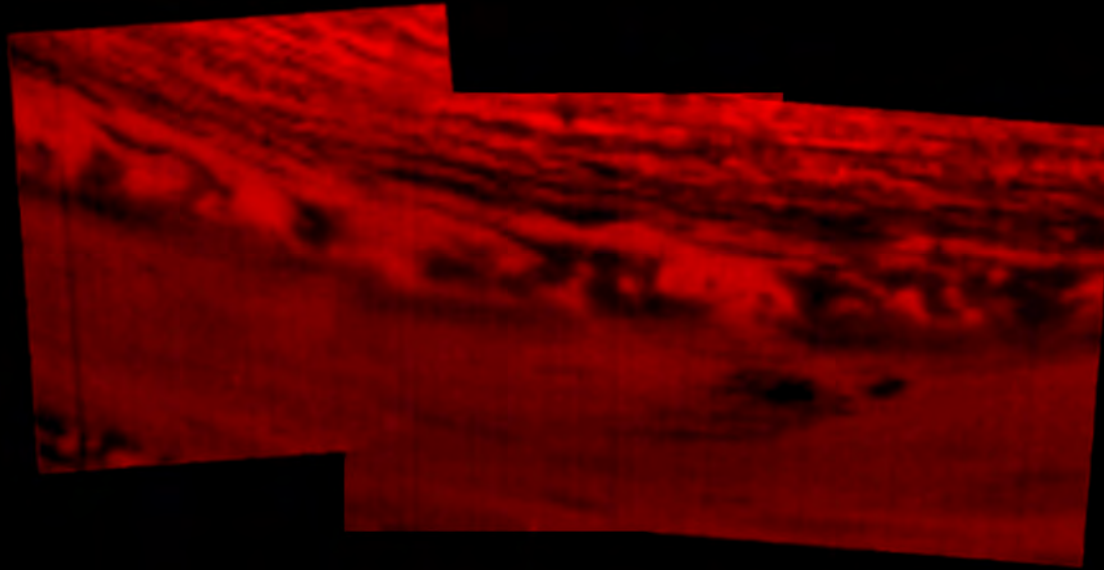
28 August 2017



Cassini ISS Observation of Impact Site
14 September 2017

VIMS Observation of Impact Site

14 September 2017



**Image taken in the thermal
IR at 5 μm . Cassini's
atmospheric entry site was
at 9.4°N 53°W.**

CIRS' Final Uploaded Command

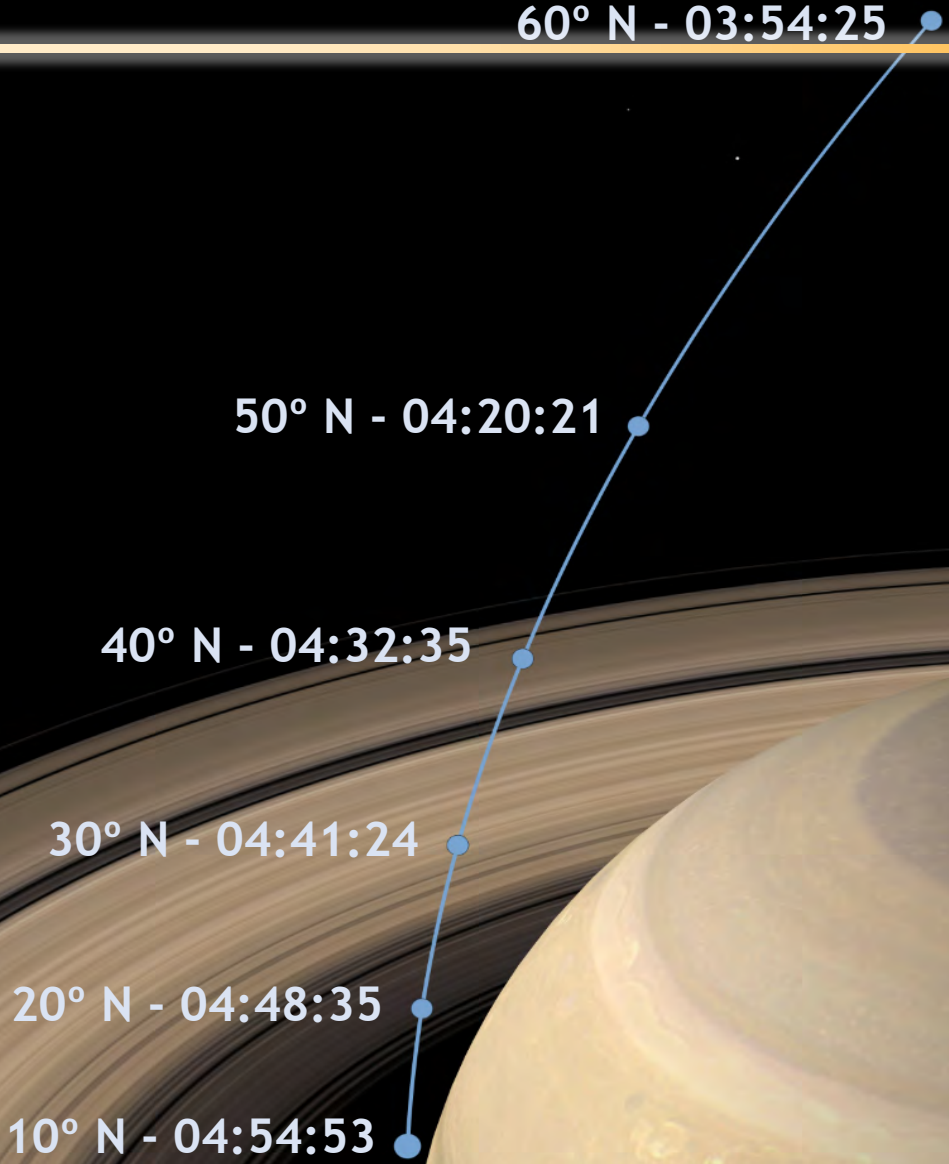
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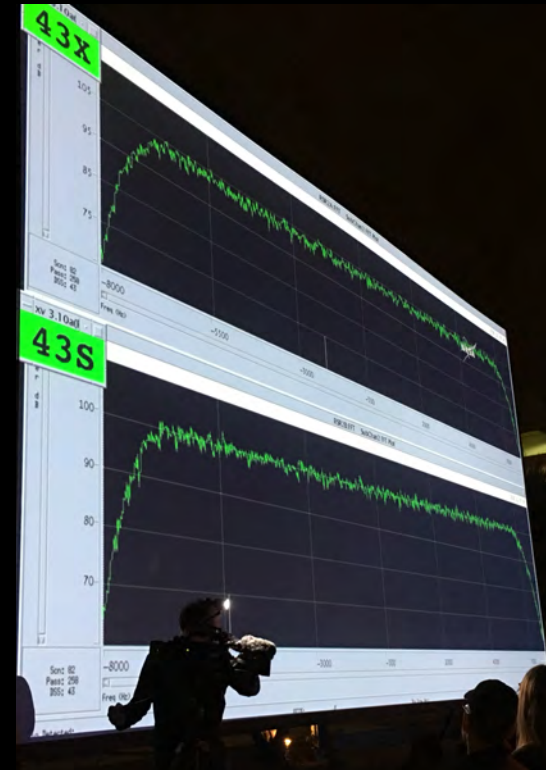
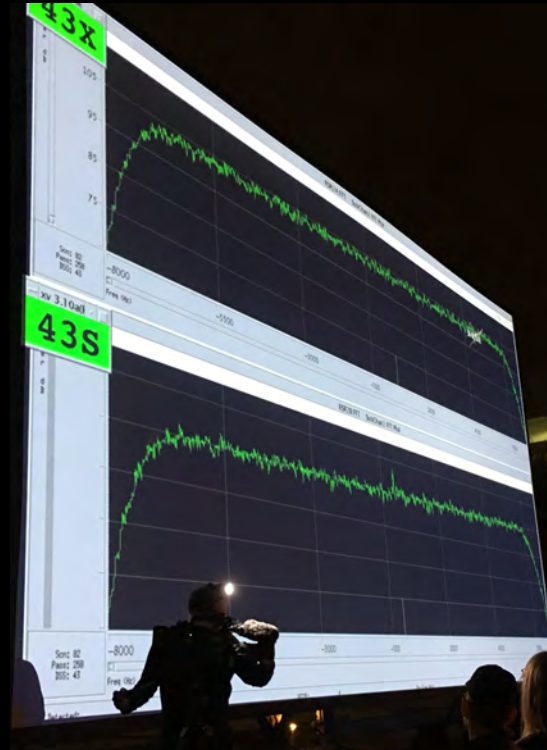
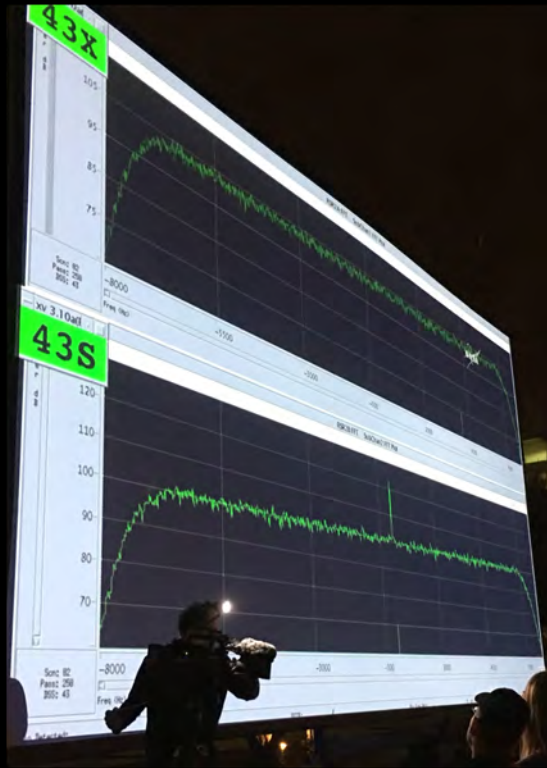
- **What it means**

Congratulations to the Cassini CIRS team for decades of tremendous achievements.

Cassini's Final Hour



Cassini's Last Communication to Earth



Cassini Composite Infrared Spectrometer: Investigation Team 1990-2017



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