## The Possibilities for Life and Human Habitability on Nearby Stars: Many Possible Homes for Our Elder(?) Race, But the Neighbors Are Likely Bacteria

- There are 100 Billion stars in our galaxy, and each has likely formed its own Earth sized planet. The Universe is 13.8 Billion yrs old, the Galaxy is about 13 Byrs old, and the solar system is 4.56 Byrs old. We live $2 / 3$ of the way out from the weak central black hole, in the suburbs. Only FGKM stars last more than 1 Byr. With Kepler, we are entering the era of precision exoplanets.
- The Earth is a highly differentiated, very unique object with an overabundance of water on the surface, a giant moon stabilizing its spin, plate tectonics keeping it warm, and shielded with a magnetic field. Life is found in an incredibly narrow range, -10 to +60 miles from the geoid surface of a body 7600 miles across. We live, work, \& die on the thinnest rind of the orange. Venus and Mars are failed harbors for life in our SS. Life has "polluted" the planet to be $\mathrm{O}_{2}$ rich, providing biomarkers.
- Life has experimented with a huge \# of forms over 4.56 Byrs on Earth, but Life is only found where liquid water exists and requires cells. The LUCA was a thermophile. Earth-life was bluegreen, squishy, and single-celled from 1.0-4.0 Byrs, and Bacteria still rule.
- Multi-cellular Life has gone through at least 4 unique stages and huge upheavals due to giant impacts. Bug Eyed Monsters, Horseshoe Crabs, or Giant Lizards would be ruling the place without them. We are the latest dominant species.

Carey Lisse, JHU Applied Physics Laboratory APS -MASPG, 25 Apr 2018

APL JOHNS HOPKINS
APPLIED PHYSICS LABORATORY

## Life Near the Solar System

## Time and Place:

## We live in the 'Burbs



REALLY In the 'Burbs....

Timescales: For the first ~ 1 Gyr, only astro- and geophysics. After, for Gyrs, the Earth harbored only Unicellular Life. The same unicellular life still dominates today, by mass \& diversity.


## The EndGame: We'll have another 5-6 Gyr of good times, and then.....

Solar Evolution Tracks For $\eta_{\odot}=0.2-0.8$


Or the Andromeda Galaxy impacts the Milky Way Galaxy 5 Gyr From Now. Or both.

## Back to Astronomy!

## Sustainable Habitable Zones :

Finding Star Systems With Regions Where Aqueous-Carbonaceous Life Could Exist On a Planetary Surface Once the System is > 1 Gyr Old
(Note We Are Ruling Out Subsurface Living
Systems, non-Carbon Systems, Polar nonAqueous Solvents like $\mathrm{NH}_{3}$, etc.)

$\triangle$ Seager S, Deming D. 2010.
R Annu. Rev. Astron. Astrophys. 48:631-72
There Are Now More Than 700 Known Planets(!) Detected by Numerous Means, and another 2700 candidates waiting to be verified.

There Are Many More to Be Found - Most Studies Are Now Pointing to about 1 Earth Sized for Each of the Stars in the Galaxy.

Exoplanet Discoveries vs. Snow-Line


None of the Known Planets is Very Close to $1 \mathrm{M}_{\text {earth }}$ at 1 AU Equivalent. (But Kepler Has Some 30 Candidates it is Watching.)
~2700 Kepler Planets : Staring in the Nörthern Cross at 1 Million Stars 24/7, Looking for Transits.




Orbital Period [days]

## Planet Sizes



Locations of Kepler Planet Candidates

Super-Earth size 1.25-2.0 Earth-size

Neptune-size 2.0-6.0 Earth-size

Giant-planet size 6.0-22 Earth-size


Kepler 62e,f: Just Announced 1.4 + 1.6 $\mathrm{R}_{\text {Earth }}$ planets around a K2V Star in 300400K Terrestrial Habitability Zone


## So Around Which Stars Do We Expect to Find Life?



Terrestrial habitability vs. stellar class and luminosity for named bright stars. Procyon A, $\alpha$ Cen A \& B, $\tau$ Ceti, $\varepsilon$ Eridani, 61 Cyg A \& B are all good candidates for life.

## Stellar Mass = Stellar Destiny




Jupiter \& Major Moons


## Trappist-1!

TRAPPIST-1 System



- "Seven temperate terrestrial planets around the nearby ultracool dwarf star TRAPPIST-1". Gillon et al., Nature 536, 437doi:10.1038/nature21360.
- Primary star is an M8V ( $\mathrm{T} \sim 2550 \mathrm{~K}, \mathrm{M} \sim 83 \mathrm{M}_{\text {jup }}$ ), 12 pc distantl L $\sim 0.0005$ Lsun, not too active
- [Fe/H] 109\% solar, > 1 Gyr old
- M_planet $=0.6-1.4$ M_earth; 3 of the planets are in the THZ; 2 are separated by just 1.6 Rearth_moon
- System has some scaled similarities to the Jovian satellite system dynamically
- Life most likely happiest at terminator, poles.


## Earth-sized planet found in nearby solar system

The three stars of Alpha Centauri


## Proxma Centauri b!

## Proxima b

- Rocky planet
- Mass: 1,3 times bigger than Earth
- Orbit: 11.2 days
- Temperature: compatible with presence of water in its liquid state
- Ultra-violet and $x$-rays 100 times stronger than on Earth
- "A Terrestrial Planet Candidate in Temperate Orbit Around Prox Cen", Anglada-Escudé et al., Nature 536, 437
- Measuring chemical abundances of Proxima is hard, use AlphaCen A \& B as proxies => solar, 6 Gyr old
- Likely no H -atmosphere, $\mathrm{M} \geq 1.3 \mathrm{M}_{\text {Earth }}$
- HST may be able to do thermal phase variations, reflected light photometry -> $R_{\text {planet }}$ (Atmo probe)
- Planet gets $\sim 0.6 \mathrm{~L}_{\text {sun, }}$, means if tidally locked subsolar pt has $\mathrm{T}=0.6^{0.25} * 1.4 \mathrm{~T}_{\text {subsolar, Earth }}$
- Life most likely happiest at terminator, poles.
- Proxima Centauri = frequent flare star, so life there needs shelter (underwater)


Age of all stars currently known to have planets within 130 light years of the Earth: Median of distribution $=3$ billion years
$=>$ It is more likely that we will find microbial life on other planets rather than animals or human equivalents


## GTAR TRUEK

So Where is the Intelligent Life in the Galaxy?
Fermi Paradox: At the speed of light, one should be able to traverse the galaxy in a few Myr. Since the Galaxy is $\sim 13$ Gyr old, and planets should have formed earlier towards the center of the galaxy, where are the neighbors?
https://en.wikipedia.org/wiki/Fermi_paradox
(Unless they're hiding. And UFO's are alien teenagers buzzing the planet.)
(See also "Alone in the Universe" by H. Smith, Am.
Sci. 99, Aug 2011)



> Photometric Imaging of 1 $1^{\text {st }}$ Detected Interstellar Object in the Solar System:

## 11/'Oumuamua

(Sept-Nov 2017 asteroid or needle-nosed spaceship?)

Bolin, Weaver,
Fernandez, Lisse+ 2017 arXiv:1711.04927v2


Figure 1: Trajectory and discovery circumstances for A/2017 Il-'Oumuamua.

Football Shaped Spheroid


## Life in the Solar System

## Necessary But Not Sufficient For Life

- Atmosphere of a Planet
- Age and Evolutionary State of A Star
- Orbital Obliquity
- Orbital Eccentcity

All Matter Too.

## Habitability Zones W/ Liquid Water Potential: Distance from Stars Where T(r) ~ 300 K (assuming $L_{*} / 4 \pi r^{2}=\sigma T^{4}$ )



(a) Surface conditions on a habitable planet within the HZ (data adapted from Kasting et al., 1993; Forget and Pierrehumbert, 1997; Kaltenegger and Selsis, 2007) \& (b) the HZ as a function of stellar type.

## Planetary Surface Temp Depends on More Than Just Distance From the Primary:

We Have \& Need a Mild Greenhouse on Earth. Venus, Twin Sister of Earth, Has a DEADLY Runaway Greenhouse!


Move Earth to Venus' Location, We Get $100^{\circ} \mathrm{F}$ Warmer, NOT $860^{\circ} \mathrm{F}$

Lammer, Selsis, et al. 2010
Astrobiology

Equilibrium temperature and albedo. The orbital distance of a planet and its albedo (the fraction of the incoming light reflected to space) determine an equilibrium temperature, $\mathrm{T}_{\text {eq }}$, corresponding to the temperature of a blackbody that reemits all it absorbs. The surface of Earth, and especially of Venus, is heated by the greenhouse effect of the atmosphere: these two planets have similar effective temperatures (the temperature associated with the radiated IR flux). However, their $\mathrm{T}_{\text {surf }}$ differ by more than $400 \mathrm{~K}!!$ Orbital distance by itself is poorly indicative of $\mathrm{T}_{\text {surf }}$ of a planet.


Boundaries of the HZ for main sequence stars of different masses, as a function of the stellar age. For stellar masses above 0.9 MSun, the main sequence lasts less than 10 Gyr. For $G$ and $F$ stars, the limits of the HZ move away from the star in response to the increase of luminosity. The evolution of stellar luminosity is taken from Baraffe et al. (1998).

## Will Uranus Be a Good Place to Live in 5 Gyr When the Sun Goes Into Its Red Giant Phase?

$\mathrm{CH}_{4}, \mathrm{NH}_{3}, \mathrm{H}_{2} \mathrm{O}$, etc.

Migrating Uranus-class exoplanets may lose their dense hydrogen atmospheres due to EUV-driven hydrodynamic escape (Kuchner, 2003; Lammer et al., 2003b, 2009a; Lecavelier des Etangs, 2007; Penz et al., 2008b). The remaining cores of these bodies may melt and outgas volatiles like $\mathrm{CH}_{4}$ and $\mathrm{NH}_{3}$, which are trapped inside the ice and may evolve to large Titan- like "terrestrial exoplanets" with reduced atmospheres.

## Life in the Solar System

Life Exists on 1 Planet in Our Solar System, and Maybe on Another and Inside the Moon of a 3rd

Pluto-Kuiper belt (short period comets)

Asteroid belt (meteorites)



Gas giants << >> Rocky planets


And life is found on the thinnest rind of our Small Blue Dot of $\sim 6370 \mathrm{~km}$ radius ......from -10 to +60 km at last reckoning.

## The Critical Roles of Water (short list :)

- Heat transfer is mediated by water on Earth
- All life we know of on Earth requires water
- Water is a major 'greenhouse' gas
- Thermostatic control (high heat capacity)
- Interacts with $\mathrm{O}_{3}$ in ultraviolet light to remove ozone
- $\mathrm{CO}_{2}+\mathrm{H}_{2} \mathrm{O}$ dissolves surface rocks




## Lessons From the Solar System:

There Is Likely Lots More to Habitability Than Just Liquid Water!
(The Laundry List of Good Happenings @ Earth :-)


Large Moon

Stabilizes tilt
Right tilt
Right tilt
Giant impacts
Giant impacts
Right amount of carbon
Atmospheric propertiespressure for life
Biological evolution

Biological evolution
Evolution of oxygen

Evolution of oxygen4
En

- $0=1$Right kind of galaxy

Right kind of galaxy

Seasons not too severe
Few sterilizing impacts after life forms
Few steriliving impacts after life forms

Enough for life BUT No runaway greenhouse effectEnough for life BUT No runaway greenhouse effect

Maintainerdegure tswum ature, composition and
Maintainterfegrare fown ature, composition and
Successful evolutionary pathway
Invention of photosynthesis
Not too much or too little
Evolves at the right time
Mature, w/ Enough heavy elements => Not small, elliptical or irregular
Right position in galaxy
Not in center, edge or halo (galactic metallicity



Mars‘ Largest Impact Basins - Evidence of Atmospheric Erosion, Too?


An Impact on the Mature Earth


Contour Map of the Moon Showing its (LHB) Impact History


## It Wasn't Always Straightforward \& Easy:

Impacts, Volcanism : Ways Into \& Out of "Snowball Earth"? The Young Sun Was Much Dimmer, ~70\% of Today's Luminosity, \& Much More Active



> Young Solar-Type Stars Are Very X-Ray Intense As They Spin Rapidly; Older Stars Slow Down \& Lose Their Violent XRay/UV Emission, Allowing Surface Life



Rotation period vs. age for the Sun-type stars of Guinan \& Ribas (2002). Solid line - power law fit to the data, indicating a spin-down with increaseing age. Rotation period vs. age for the Sun-type stars of Guinan \& Ribas (2002). Solid line - power law fit to the data, indicating a spin-down with increasing age. Luminosity ${ }_{\mathrm{x}-\mathrm{ray}} \sim 1 /$ Period $^{3}$


Figure 12. Integrated surface fluxes of C III $977 \AA, \mathrm{O}$ vi $1032 \AA$, and Fe XVIII $975 \AA$ for some of our FUSE targets as a function of rotation period. The straight lines are power law fits to the measured data. The coefficients of the three fits are printed at the right end of the panel.


## Small Eccentricity, Obliquity, Stable Orbits:

## The Unsung Heroes

- Think of How Much We Like the Temperate, EvenedOut Tropical Climate
- And How Much the Winter/Summer Cycle Stresses and Forces Our Biota
- Moon Stabilizes Our Tilt (Obliquity)
- Mars has no such Stabilizer, and Its Polar Tilt Has Varied Enormously on $\leq$ Myr timescales (imagine London moving to the North pole and back)


## Example of A Failed Biosphere: Mars

Frost $\left(\mathrm{CO}_{2}\right)$ seen by Viking 2 Lander (ca. 1978)

## Mars, Relative to Earth

## Smaller size:

Faster Cooling > Mostly Solid Core > Very Weak Magnetic Field Lower Gravity > Loss Of Atmosphere To Space
No Tectonics > Limited Chemical Cycling Of Elements
Location:
Less Solar Irradiance
Higher Eccentricity Of Orbit > Less Stable Long-term Climate


Question: What if Venus had formed where Mars is and Mars Had fromed where Venus is?

Would Venus then be a cool Earth twin and Mars have stayed Warm and wet? Or would Mars have just lost its water via Hstrippping but kept warm enough to keep tectonics going?

Timescales: For the first ~ 1 Gyr, only astro- and geophysics. After, for Gyrs, the Earth harbored only Unicellular Life. The same unicellular life still dominates today, by mass \& diversity.



Southern highlands

Olympus Mons
Highs (red, white) and Lows (blue) of
Spirit
Mars

Northern
lowlands

Opportunity

Did Mars Once Have a Huge Northern Ocean?

## Life On Earth Over Time:

We Weren't The First or Second Animal Life Form Chosen to Run the Planet We're Batting Cleanup (maybe).

Intelligent Life's Duty Cycle on Planet Earth is Only 4 Myr/4.56Gyr < $\mathbf{1 0}^{-3}$ !!

## Milestones in the History of Life



OTH, unicellular life, which dominates Earth's biomass and its history, seems quite likely elsewhere! We see Archaea/Bacteria in extreme environments: high or low temperatures, high acidity, deep underground, rich in dioxins or arsenic or sulfur, can be agreeable environments for extremophiles! Pretty much wherever there is exploitable Free Energy, carbon compounds, \& some liquid water....in arsenic rich lakes, eating dioxins, 10 miles down in rock or way under the Antarctic ice cap. One 30



Acid, heat loving thermophiles in Yellowstone.

## The Big Tree of Life



You Are Here


All life is cellular...defeating entropy by a steadystate burning of nutrients in an $\mathrm{H}_{2} \mathrm{O}$ medium


- Confined within the cell membrane
- Interacts with its environment
- Complex biochemistry in aqueous solution
- All functions controlled by genetic information


Hypothesis for the formation of cellular life on Earth from $\mathrm{FeS}_{x}$ and $\mathrm{NiS}_{x}$ cell-like precipitates. The steady state flow of "nutrients" is provided by movement of hot, alkaline hydrothermal solutions associated with magma into colder, more acidic Hadean ocean water. (After Hall \& Russell 2003).

If this picture is correct (e.g. recent findings of bacterial mats in the Marianas Trench, Cameron et al. 2012 AGU), then life in subsurface, internally heated oceans (e.g. Europa, Enceladus, ancient Mars N. hemisphere) is entirely plausible!!


1 million microorganisms/ml of seawater (upper ocean)

1 billion microorganisms/g of soil

## $5 \times 10^{30}$ microorganisms on Earth

## $10^{31}$ viruses

~ $10^{11}$ stars in our Galaxy


Less than $1 \%$ can be cultivated!

## Estimated number of all distinct species, not including bacteria


*The ocean database was substantially more complete than the database for the entire Earth,
so PLaS Biology used the oceun database to estimate the total number of species in this taxon.

## Diversity of Microbial Life

(aka bad breath, stinky feet, garbage, bread mold, pond scum, shower mold, poo...)



UniCellular Life Ruled Then...and Now

(On us, in us, on the ground, in our food, in the air)

~1 trillion cells in human body 10 times more microorganisms in our body


30,000 genes in human genome 100 times more microbial genes in our body


Multicellular Life Has Tried Many Different Ways to Occupy the Planet's Biomes



## Something Has Infrequently Caused Large, Sudden Drops In Biodiversity on Our Planet

Periodic extiction events for past $\mathbf{2 5 0} \mathbf{M}$-years.
The straight line fits $\mathbf{2 6}$ M-year-peiodicity



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Is this an alien? Probably not, but of all the animals on Earth, the tardigrade might be the best candidate. That's because tardigrades are known to be able to go for decades without food or water, to survive temperatures from near absolute zero to well above the boiling point of water, to survive pressures from near zero to well above that on ocean floors, and to survive direct exposure to dangerous radiations. The far-ranging survivability of these extremophiles was tested in 2011 outside an orbiting space shuttle. Tardigrades are so durable partly because they can repair their own DNA and reduce their body water content to a few percent. Some of these miniature water-bears almost became extraterrestrials recently when they were launched toward to the Martian moon Phobos on board the Russian mission Fobos-Grunt, but stayed terrestrial when a rocket failed and the capsule remained in Earth orbit. Tardigrades are more common than humans across most of the Earth. Above: a mm-long tardigrade crawls on moss.

## BioMarkers: Earth's Atmosphere

> We live in a very special atmosphere Not too Cold (Mars, Moon)
> Not too Hot (Venus, Mercury)
> Not too Dense (Venus)
> Not too Sparse (Mars, Mercury, Moon)

Earth's Atmosphere has been polluted!
The "typical" $\mathrm{CO}_{2} / \mathrm{N}_{2} / \mathrm{H}_{2} \mathrm{O}$ rocky planet atmosphere (Venus, Mars) has been forced out of equilibrium by Plants.
Turn off today's plants, and rocks will reabsorb the $\mathrm{O}_{2}$. (=> Be kind to your chlorophyll laden friends.)

Mild Greenhouse + Forced $\mathrm{O}_{2}+$ Solid/Liquid/Gaseous Water $+\mathrm{CO}_{2}$ Buffer $=$ Life $?$


Figure 3. A sketch showing the expected evolution of the Earth's atmosphere. From the NASA website (http://science.nasa.gov/ headlines/y2002/10jan_exo-atmospheres.htm?list161037).

## The Deep Impact/EPOXI Spacecraft Stared at the Earth \& Mars as it Traveled from Comet to Comet



EPOXI Views of the Earth.
"Visible" = RGB Colors.
"NIR-R" = Chlorophyll Red Edge Color.

Top: Eastern Hemisphere, Africa Prominent.

Middle: Western Hemisphere, Centered on the Pacific Ocean.

Bottom: Rotationally Averaged Image, Showing What a Realistically Imaged Exo-Earth Will Look Like.


## EPOXI Observed the Earth as It Rotated 1 Full Cycle, and as the moon Passed In Front



Best Measured EPOXI Optical/Near-IR Bio-Signatures Were Atmospheric Gases: $\mathrm{H}_{2} \mathrm{O}, \mathrm{O}_{2}, \mathrm{CO}_{2}, \mathrm{CH}_{4}, \mathrm{~N}_{2} \mathrm{O}$, Not Chlorophyll "Red Edge" at $\sim 850$ nm


## Earth's Atmospheric Signatures Farther in the Near-Infrared



Earth's
Atmosphere Stands Out in Simple Color Photometry, Too.

Color - Color Plot of Solar System Worlds


## Life \& Out of Equilibrium Atmospheres

- Critical Idea: Our Current Atmosphere is Very OxygenRich and Carbon Dioxide Poor.
- If We Turn Off All Green Plants Tomorrow, the Earth's Rocks Would Re-Absorb the Free $\mathrm{O}_{2}$ and Return Us to a "Normal" Rocky Terrestrial Atmosphere
- This is What Makes $\mathrm{O}_{2}$ (and $\mathrm{O}_{3}$, a UV photolytic product of $\mathrm{O}_{2}$ ) Such Good BioMarkers for Chlorophyll-based Life

Earth has
$\mathrm{CO}_{2} \& \mathrm{~N}_{2}$ like Mars \& Venus, but also Unique $\mathrm{O}_{2}, \mathrm{O}_{3}$, $\mathrm{H}_{2} \mathrm{O}$ \& a Thermal Inversion.








## Net Photosynthesic Reactions Driving Alterations of the Earth's Atmosphere Today: Destruction of Native $\mathrm{CO}_{2}$ and Creation of $\mathrm{O}_{2}\left(\mathrm{O}_{3}\right)$



## Alone in the Universe

# Despite the growing catalog of extrasolar planets, data so far do not alter estimates that we are effectively on our own 

Howard A. Smith American Scientist 99, Aug 2011

TThe first known extrasolar planet (that is, a planet orbiting around a normal star other than our Sun) was 51 Pegasi b, discovered only about 16 years ago. Today, thanks largely to NASA's orbiting Kepler satellite, there are more than 1,500 known extrasolar planets (or planetary candidates), about 431 with confirmed detections and reasonably welldetermined parameters (such as radius, mass and orbital characteristics).
These discoveries are exciting not only to astrophysicists; the public is also keen to learn about Earthlike extrasolar planets and the possibility that some might host life, even intelligent life. Last year the Royal Society of London sponsored a symposium with the dramatic title "The Detection of Extra-terrestrial Life and the Consequences for Science and Society." Participants observed, "Should it turn out that we are not alone in the universe, it will fundamentally affect how humanity understands itself," and cited polls suggesting that most people believe we do have cosmic company (half of this group think aliens have already visited).

The public wants to believe in aliens (or extraterrestrial intelligence, ETI), some say, because they believe that "EII comes from 'utopian societies which are free of war, death, disease or any other ... mid20th century problems' and could 'help mankind overcome its problems.'"

Scientists and the press often encourage people to adopt these sensational attitudes. Last March, for example, the New York Times ran an op-ed piece on extrasolar planets by astronomer Ray Jayawardhana under an image of a sky full of stars labeled with "THEM" signposts. The author was reassuring: People should not fret about "life elsewhere, especially if it turns out to be in possession of incredible technology [that can] make us feel small and insignificant." He did not speculate on the possibility that this science-fiction scenario might instead make people feel inconsequential or prompt them to treat each other, or our planet, casually.

The dawn of the era of extrasolarplanet (or exoplanet) discovery provides us with the first bits of hard information to use in reconsidering the likelihood of this popular attitude. The snin has heen that hecanse nlanets
ligent beings are the inevitable product of life and evolution. This view has been the traditional attitude, and one typical portrayer was Percival Lowell, an astronomer famous for his advocation of canals on Mars, who wrote in his 1908 book, Mars as the Abode of Life:

From all we have learned of its constitution on the one hand or of its distribution on the other we know life to be as inevitable a phase of planetary evolution as is quartz or feldspar or nitrogenous soil. Each and all of them are only manifestations of chemical affinity.
Today we know that Mars has no artificial canals, and that this assertion was unsupported wishful thinking. Don Goldsmith and Tobias Owen, in their classic book, The Search for Life in the Universe (1993 edition), present a more modern view:

We anticipate that all planetary systems will have a set of rocky inner planets, with atmospheres produced by outgassing, weathering and escape, for the same reasons

Assuming a radius of 1000 pc, or 100 human generations, as Search Space.

## The Possibilities for Life and Human Habitability on Nearby Stars: Many Possible Homes for Our Elder(?) Race, But the Neighbors Are Likely Bacteria

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- Multi-cellular Life has gone through at least 4 unique stages and huge upheavals due to giant impacts. Bug Eyed Monsters, Pill Bugs, or Giant Lizards would be ruling the place without them. We are the latest dominant species.

Carey Lisse, APL SD/SRE
APL, Laurel, MD 26 April 2013

